Statistical study of electron bunching in auroral Langmuir waves

Micah P. Dombrowski¹, Jim LaBelle¹, Craig A. Kletzing², Scott R. Bounds², Iver H. Cairns³, Stephen R. Kaeppler⁴

¹Department of Physics and Astronomy, Dartmouth College, Hanover, NH, USA.
 ²Department of Physics and Astronomy, University of Iowa, Iowa City, IA, USA.
 ³School of Physics, University of Sydney, Sydney, Australia.
 ⁴Department of Physics and Astronomy, Clemson University, Clemson, SC, USA.

9 Key Points:

1

2

3

10	•	The CHARM-II sounding rocket observed 57 correlations between electrons and Lang-
11		muir wave phases.
12	•	Correlations reveal a specific relationship between electron beam flux gradient and energy
13		flow.
14	•	Simulation shows time-of-flight effects on the particle distribution are a plausible expla-
15		nation.

 $Corresponding \ author: \ Micah \ P. \ Dombrowski, \ \texttt{micah.p.dombrowski} \\ \texttt{Corresponding author: Micah } P. \ Dombrowski \\ \texttt{Micah } P. \ Dombrowski \\ \texttt{Corresponding author: Micah } P. \ Dombrowski \\ \texttt{Corresponding author: Micah } P. \ Dombrowski \\ \texttt{Corresponding author: Micah } P. \ Dombrowski \\ \texttt{Micah } P. \ Dombrowski \\ \texttt{Corresponding author: Micah } P. \ Dombrowski \\ \texttt{Corresponding author: Micah } P. \ Dombrowski \\ \texttt{Micah } P. \ Dombrowki \\ \texttt{Micah } P. \ Dombrowski \\ \texttt{$

16 Abstract

Langmuir waves are a fundamental consequence of electron beam-plasma interactions in space 17 physics. Wave particle correlator experiments have proven to be an effective way to probe the 18 physics of Langmuir wave growth, damping, and particle trapping, but previous experiments 19 were limited mainly to special cases. The CHARM-II rocket flown into active nighttime aurora 20 included a correlator and encountered Langmuir waves along much of its trajectory. Careful 21 consideration of the data yields 57 events in which significant correlations were detected be-22 tween electrons and Langmuir wave phases. The majority of the selected correlations indicated 23 energy exchange slightly dominant over particle trapping. Examination of the component of 24 the correlation resulting from energy exchange reveals a striking relationship between electron 25 beam dynamics and the nature of the wave-particle correlation: whenever the beam flux at 26 the measured electron energy was increasing with time, the phase of the resistive component 27 of the electron bunching implied energy transfer from the beam particles to the wave field, and 28 when the electron beam flux was decreasing, the reverse occurred. Two possible explanations 29 for this effect are the time variation in the interaction of a given energy with the waves in a 30 beam which is time-varying due to dispersion effects, and short-term enhancements in wave 31 growth or damping on a time-varying beam due to time-of-flight effects on the particle dis-32 tribution. The latter mechanism is investigated and found plausible using a numerical test 33 particle simulation. 34

35 1 Introduction

Langmuir waves, also known as electron plasma waves, are one of the most fundamental 36 properties of a plasma, having been first observed in discharge plasmas in the early days of 37 plasma physics (Langmuir, 1928). They result from the interaction of electron beams with 38 plasmas and are ubiquitous in space plasma, including, as examples: the solar wind, where 39 they generate radio bursts (Lin et al., 1981); planetary foreshocks (Filbert & Kellogg, 1979; 40 Gurnett et al., 1981); and the auroral ionosphere (Kintner et al., 1995; M. H. Boehm, 1987; 41 McAdams, 1999; Samara, 2005), where they mediate energy transfer between the beam and 42 thermal plasmas. Langmuir waves can generate nonlinear structures of fundamental interest to 43 plasma physics, as well as linear eigenmode effects in inhomogeneous plasmas (McAdams et al., 44 2000; Ergun et al., 2008). Due to their significance and abundance in the space environment, 45 Langmuir waves are a subject of intensive study, particularly with regard to their eigenmode 46 structures (Malaspina et al., 2012; Graham & Cairns, 2013), nonlinear processes (Graham 47 & Cairns, 2014), three-dimensional effects (Malaspina & Ergun, 2008; Dombrowski et al., 48 2012), and wave-particle correlations (Ergun, Carlson, McFadden, TonThat, & Clemmons, 49 1991; Ergun, Carlson, McFadden, & Clemmons, 1991; Muschietti et al., 1994; Kletzing et al., 50 2005; Kletzing & Muschietti, 2006). 51

Dedicated rocket- or satellite-borne wave-particle correlator experiments have proven to 52 be an effective way to probe the interactions between Langmuir waves and electrons in space 53 plasmas. A detailed summary theory of expected theoretical results from such instruments 54 is given by Kletzing and Muschietti (2006). The phase bunching of the electrons in the field 55 of the wave can be considered as a superposition of two components, a 'resistive' component 56 which is in phase with the wave electric field and represents energy transfer either from wave 57 to particles or vice versa, and a 'reactive' component which is in quadrature phase with the 58 wave field and is a signature of electrons trapped in the wave. The width of the resonant 59 part of the electron energy spectrum is inversely related to the length of the wave packet, and 60 the resonant component is unipolar, whereas the reactive component is bi-polar; that is, the 61 reactive component is negative over part of the energy range and positive over the other part. 62 This signature implies that if the energy resolution is too coarse, the reactive component will 63 not be detected at all because its positive and negative parts will cancel out. 64

An early version of wave-particle correlator flown on a sounding rocket in auroral plasma detected a strong correlation between beam electrons and Langmuir/upper hybrid wave electric

fields over an interval of several hundred seconds (Gough et al., 1990). Ergun et al. (1998) 67 flew a wave-particle correlator on an auroral rocket which worked by sorting detected electrons 68 according to the quadrant of the Langmuir wave phase. This experiment revealed evidence 69 of wave-particle interactions but did not address resistive versus reactive components. More 70 recently, Kletzing et al. (2005) reported results from a wave-particle correlator with very good 71 energy resolution (10%) and phase resolution (22.5 degrees). Launched into nighttime aurora, 72 the experiment measured correlations associated with a relatively small number of the most 73 intense Langmuir waves encountered, and the results gave a strong indication of wave trapping 74 of the bunched electrons in those examples. A wave particle correlation experiment was flown 75 on the Freja spacecraft but was not used much in flight (M. Boehm et al., 1994). 76

The success of these previous wave-particle correlator experiments inspired subsequent 77 rocket experiments using a correlator very similar to that of Kletzing et al. (2005). The 78 CHARM-II sounding rocket (NASA mission 40.025) was launched from the Poker Flat Research 79 Range near Fairbanks, AK, at 9:49 UT/22:46 MLT on 16 February 2010, reaching an apogee 80 of 802 km. The launch was into an active substorm expansion phase, characterized by a 20 nT 81 bay in the H-component of the magnetic field observed by GOES 11. The payload carried a 82 Dartmouth High-Frequency Experiment and University of Iowa Correlator, as well as a number 83 of other primary and contextual instruments. One particularly intense event encountered was 84 reported by Kletzing et al. (2016). 85

The CHARM-II correlator measured over fifty examples of significant Langmuir wave-86 electron correlations—over an order of magnitude more than all previous experiments com-87 bined. The large number of examples for the first time reveals a previously unobservable pat-88 tern relating the phase of the correlation to the temporal or spatial gradient in the causative 89 90 beam electrons. Section 2 describes the instruments which make up the Correlator system and the form of the data. Section 3 describes the method used to identify significant correlation 91 events, and then presents several case study events from the data set as well as statistical 92 observations regarding the data set as whole. Section 4 summarizes these results and analyzes 93 two models developed to explain the observations. A numerical test-particle simulation used 94 to test the models is described in Appendix Appendix A. 95

96 **2** Instrumentation

Accurate, in-situ correlation of Langmuir waves and electrons requires three primary pieces: a wave instrument covering the range of frequencies in which Langmuir waves are expected, high-speed particle detectors covering the range of potentially resonant energies with high resolution, and the correlation hardware itself, which processes these data streams and returns the desired statistics.

The Dartmouth High-Frequency Experiment (HFE) detects the potential difference be-102 tween two 2.5 cm spherical probes, separated by 30 cm along a fixed, axial boom parallel to 103 the payload's spin axis, which is maintained within ten degrees of the ambient magnetic field 104 by an attitude control system (ACS). Therefore, this potential difference provides an estimate 105 of the the parallel component of the electric-field, exactly the desired component for maximum 106 sensitivity to Langmuir waves near the plasma frequency, which have electric fields predom-107 inantly parallel to the background magnetic field. Active preamplifiers inside each spherical 108 probe assure that the antenna functions as a double-probe over the 0-5 MHz frequency range. 109 The signal is band-pass filtered to a 100 kHz to 5 MHz band, and regulated by an Automatic 110 Gain Control (AGC) system to optimize usage of the dynamic range. The regulated HF sig-111 nal directly modulates a 5 MHz-bandwidth S-band transmitter, and the resulting waveform 112 is continuously digitized in the ground telemetry station at 10 MHz, with 12-bit resolution. 113 The AGC control signal is digitized onboard at 20 kHz and telemetered on a separate digital 114 telemetry link. This instrument is the latest iteration of a design which has flown on numerous 115 other rocket campaigns in both E_{\parallel} and E_{\perp} configurations, including HIBAR (Samara et al., 116

2004), PHAZE II (McAdams et al., 1998), SIERRA and RACE (Samara & LaBelle, 2006a,
2006b), and ACES (Kaeppler et al., 2013).

For detection of electrons associated with the Langmuir waves, CHARM-II used eight 119 'bagel' particle detectors provided by the University of Iowa. These detectors, named for the 120 bakery items they resemble, have energy acceptance ranges of 10% and are characterized by a 121 large geometric factor, as required to provide good counting statistics for correlation with high-122 frequency waves. The detectors are aligned with the rocket's spin axis, with a 10°-wide field of 123 view, and thus are always observing the field-aligned component of incoming particles when the 124 payload is aligned to the magnetic field. The bagel detectors are tuned to logarithmically spaced 125 energy levels ranging from 200 to 1050 eV. The resulting approximately 10% energy resolution 126 is significantly better than achieved in early correlator experiments, and has been shown in 127 recent experiments to be adequate for detecting both the resistive and reactive components of 128 wave-particle correlations in nighttime aurora (Kletzing et al., 2005). 129

The University of Iowa Wave-Particle Correlator combines the above described Langmuir 130 wave and electron measurements to measure wave-particle correlations. The correlator was 131 previously flown on two rocket missions, RACE and CHARM, and is described in detail by 132 Kletzing et al. (2016). The correlator takes an input waveform from the HFE and uses it to 133 control a phase-locked loop (PLL) circuit running at 16 times the frequency, locked to the 134 frequency of the highest-amplitude component of the incoming wave. When Langmuir waves 135 are unstable, they dominate the HFE wave signal, and under this condition the PLL produces 136 a clean square-wave locked in phase to the Langmuir wave, but at sixteen times its frequency. 137 This square wave can then be used to sort incoming electron counts from each of four detectors 138 into phase bins which accumulate during an integration period—1 ms in the CHARM-II case, 139 corresponding to hundreds of wave periods per timeslice. CHARM-II included two correlators, 140 each analyzing electron counts from four of the eight 'bagel' detectors. 141

CHARM-II also included an Electrostatic Electron Pitch Angle Analyzer (EEPAA), a 'top-hat' style detector which counts electrons with a 50 ms integration time, sorted into 15°-wide pitch angle bins and 47 logarithmically spaced energy bins from 15 eV to 15.5 keV. Flux-gate magnetometers, Langmuir probes, and DC/ELF/VLF electric field and wave measurements rounded out the instrument complement.

¹⁴⁷ **3 Data Presentation**

Figure 1 shows a survey of wave and electron spectra from the active period of the 148 CHARM-II flight, 325-800 s after launch. The top panel is an energy spectrogram of 15° and 149 30° pitch angle EEPAA data, the middle panel is a 300-900 kHz spectrogram of HFE data, and 150 the bottom panel is a plot of \log_{10} of total counts detected by the eight Bagels. In the EEPAA 151 data, an inverted-V structure is clear from approximately 610 to 660 s, with a more tenuous 152 one from 500 to 560 s. In the HFE data, the upper cutoff to noise which is near 500 kHz at 350 153 s but rises to 800 kHz after 700 s is interpreted as the plasma frequency ω_p , which acts as an 154 upper bound to whistler modes in 'underdense' plasma ($\omega_p < \omega_c$). In the auroral ionosphere 155 ω_c typically exceeds 1 MHz, and the plasma frequency is clearly much lower than the upper 156 hybrid frequency (~ 1.4 MHz), and therefore easily identified despite slight (< 10 degrees) 157 variations in the rocket's alignment parallel to the magnetic field. For resonant electrons in the 158 energy range of the 'bagel' detectors, it can also be deduced that the Langmuir wavelengths 159 will be tens of meters. There are many instances visible in Figure 1 where increased particle 160 counts are accompanied by wave activity near ω_p . The activity near 650 to 660s corresponded 161 to a particularly powerful event in which the peak electric field, estimated to be 1-3 V/m, 162 momentarily saturated the on-board DC electric field instrument. 163

The correlator system returned 490 seconds of valid data, providing a matrix with accumulated counts s(t, p, E), where t labels each of 490,000 1-ms timeslices, p (=0...15) labels each of the 16 phase bins, and E (=1...8) labels each of the 8 energy levels. The interpretation of



Figure 1. A summary plot of the active period of the CHARM-II flight. Both time after launch and altitude are shown on the x-axis. Top is a spectrogram of 15° and 30° pitch angle EEPAA particle data, with energy on the y-axis, and \log_{10} of the differential energy flux [eV/(cm² * ster * s * eV)] as color intensity. Middle is an HFE spectrogram, with frequency on the y-axis, and color following wave power in decibels. Bottom is a plot of \log_{10} of the total counts among all eight Bagels detectors. Inverted-V structures occur at 610-660 s and 500-560 s, the latter tenuous. The upper cutoff to noise in the HFE panel is interpreted as the plasma frequency. The electron density increases dramatically after 700 s, and broadband noise signals, presumably whistler-mode auroral hiss, occurs at frequencies below f_{pe} .

the phase bin, in terms of the phase between the electrons and the input waveform, varies with the frequency due to phase shifts in the electronics and other delay sources, all of which were carefully calibrated pre-flight. Therefore, a first step taken to aid comparability of timeslices is to shift all phase bins to the same baseline, based on the recorded correlator frequency.

While the most direct way of looking at this data might be to display raw counts vs. phase and time, a more effective quantitative analysis is achieved by examining the Poisson z-score for the data,

$$\sigma(t, p, E) = \frac{s(t, p, E) - \bar{s}(t, E)}{\sqrt{\bar{s}(t, E)}},\tag{1}$$

where \bar{s} is the mean particle count of the timeslice, $\bar{s} = \frac{1}{16} \Sigma_p s(t, p, E)$. This parameter is 174 a measure of particle over- or under-density in each phase bin with respect to the expected 175 value, assuming all phases are equally probable. The left panel of Figure 2 displays example 176 results of this analysis, showing z-scores expressed as a color scale, with white being zero, red 177 expressing negative values (deficit of counts relative to expected value) and blue expressing 178 positive values (surfeit of counts relative to expected value), for 101 timeslices (on the x-axis), 179 as a function of the electric field wave phase (on the y-axis). As mentioned above, calibration 180 data have been applied to shift the phase bins from the raw values so that they reflect phase 181 relative to the electric field wave. 182

Figure 2 shows a high- $|\sigma|$ event lasting for multiple timeslices in the middle of the plot 183 (near t = 50 ms). However, for most of the times and phase bins the z-scores are insignificant, 184 appearing as 'salt and pepper' in the plot. Cursory examination of the entire data set reveals 185 that most timeslices can be discarded due to a lack of significance. Furthermore, many features 186 that appear significant turn out to be results of natural and instrumental interference. For 187 example, for each timeslice the correlator records whether the PLL is 'locked' or not, and to 188 what frequency it is locked. Less than 15% of the data set has both correlators locked. In addi-189 tion, a timeslice cannot be considered reliable merely from the presence of a lock state, as this 190 says nothing about the presence of interesting activity at that time, or whether the correlator 191 is locked to the appropriate Langmuir frequency (which can be determined by comparing the 192 'lock' frequency to the HFE data). With so much data, manual inspection was not a practical 193 or desirable method to identify reliable lock or significance, so an automated algorithm for 194 event identification was developed. 195

The reduction analysis is motivated by an expectation, based on theory (Kletzing & Muschietti, 2006), of how significant wave-particle correlations will manifest themselves in the data: as a sine wave in the phase bins, with a quarter of the bins having a statistical excess of counts, and a quarter having a deficit. The pattern arises because of the bunching of the particles in the electric field of the Langmuir waves to which the PLL is locked.

The first step towards automated reduction of the correlator data to identify discrete events—such as that shown in the middle ($t \sim 50$ ms) of Figure 2—is to re-bin the 16 phase bins into four reduced phase bins. Figure 3 shows the scheme: the 16 phase bins p (orange) are summed into four reduced phase bins p' (color coded by p'),

$$s'(t, p', E) = \sum_{n=0\dots3} s(t, 4p' + n, E),$$
(2)

in order to emphasize the expected pattern of a quarter of the bins having overdense and another quarter having underdense counts. The re-binning was repeated four times (q = 1...4, depicted as individual columns of p' bins), shifted by one raw bin for each, to cover all possible sinusoidal patterns that might result from a wave-particle correlation event,

$$s'_{q}(t,p',E) = \sum_{n=0\dots3} s(t,4p'+n+q,E).$$
(3)

A z-score,

$$\sigma'_{q}(t, p', E) = \frac{s'(t, p', E) - \bar{s}}{\sqrt{\bar{s}}},$$
(4)



Figure 2. An example plot showing (left) the Poisson z-scores (σ) vs. phase and time for 101 ms of CHARM-II Correlator data. The z-score from -5 to 5 is shown with the blue-to-red color scale, electric field phase is on the y-axis, and time (in relative ms) on the x-axis. A high- $|\sigma|$ event lasting for multiple timeslices occurs in the middle of the plot, but for many of the timeslices, the z-scores are insignificant. In the left line plot the σ values for the central timeslice are shown, along with the base (unity amplitude) functions for an I/Q Resistive/Reactive fit, while in the right line plot they are shown with the fit amplitudes and their sum. The fit is reasonable, as shown by the r^2 goodness-of-fit value. In this case the Q component dominates, implying electrons trapped in the wave. Only 15 of the 57 events were of this type.



Figure 3. A map of the rebinning done at each timeslice to the raw correlator counts. Each large group is a time integration (singlet, doublets, triplet), and each of those contains four shifted rebinnings from the original 16 correlator bins of s(t, p, E) (shown in dark orange), to the 4 bins of $s'_q(t, p', E)$ (shown in other colors, grouped by p').

²¹⁰ was then calculated for each of these reduced bins.

To account for events which span multiple timeslices, two 'doublet' sets and one 'triplet' set were constructed at each timeslice, acting as additional arrays of timeslices (the large divisions in Figure 3). For these, the means and the counts in each reduced phase bin were integrated in time over the two or three raw timeslices, with the doublet sets defined as

$$s'_{q,t2b}(t,p',E) = \sum_{\tau=t-1,t} s'_q(t,p',E) \text{ and } s'_{q,t2f}(t,p',E) = \sum_{\tau=t,t+1} s'_q(t,p',E),$$
(5)

²¹⁵ and the triplet

$$s'_{q,t3}(t,p',E) = \sum_{\tau=t-1,t,t+1} s'_q(t,p',E),$$
(6)

i.e. the doublets integrate either the timeslice prior or after, and the triplet both. All three sets then have their associated $\sigma'_{q,t2b}$, $\sigma'_{q,t2f}$, and $\sigma'_{q,t3}$. Thus, the process yields four total σ' arrays over the four q values, times four timeslice arrays (singlet, doublets, triplet), or sixteen total arrays.

Finally, as a criterion to identify timeslices with interesting events, we find the global minimum and maximum σ' over the sixteen arrays at each timeslice. The difference between the minimum and maximum, Δ , provides a scalar measure of how well a given timeslice matches the expected signature of a wave-particle correlation event.

Initially, a simple global Δ threshold value was used to identify events, resulting in hundreds of candidate events; however, thorough investigation of these candidates revealed many

false positives, including many events which were disqualified after examination of diagnos-226 tic and contextual data. Figure 4 shows the two major factors which lead to disqualifying 227 candidate events. Panels (a) and (b) contain an energy spectrogram derived from the bagel 228 detectors, and a line plot of the logarithm of counts from the selected bagel, with an 'X' at 229 bottom indicating when the correlator is locked. The spectrogram in panel (c) shows HFE 230 data concentrated around the plasma frequency—about 530 kHz in this case—with a white 231 line indicating the frequency to which the correlator was locked. The overlaid dashed black 232 line indicates the hand-estimated Langmuir frequency at the center time. Note that the signal 233 at this estimated frequency is relatively weak, short-lived, and surrounded by noise. The final 234 two panels show raw HFE waveforms on two different timescales, with the shorter timescale 235 in panel (e) revealing a powerful interfering spike-like signal at a cadence of about 15 μ s. This 236 signal, originating on the rocket payload from an unknown source, is frequently seen in the 237 HFE data, and in some cases is the highest-amplitude component of the waveform. 238

Either factor—weak signal at the Langmuir cutoff or obvious payload interference would reduce confidence in the correlator phase binning. While in principle the correlator hardware filters out low-frequency interference, here it combines with a poorly defined Langmuir cutoff to yield a highly unstable correlator lock frequency, fluctuating between unrealistically high and low values. In panel (b) we see that the hardware's lock status is intermittent around the center time, definitively indicating a lack of significant coherence in the data. Together, these indicators motivate the discarding of this event and many others like it.

The large number of false identifications in the initial run also revealed that the Δ 246 threshold needs to be different for each bagel detector. With these issues in mind, the event 247 identification algorithm was altered to iteratively optimize the Δ threshold for each bagel. 248 249 Subsequently, manual screening was applied based on considerations such as those illustrated in Figure 4. Figure 5 shows an overview of the final set of 57 thresholded, hand-screened events, 250 as a scatter plot versus bagel number and time, with a histogram versus time overlaid. The 251 final set of events is tabulated and plotted in the Supporting Information. From the histogram 252 it is clear that a majority of the selected correlator events occur during times of high flux 253 at lower energies coincident with the boundaries of inverted-V structures at higher energies, 254 particularly the tight cluster near 652 seconds coincident with the end of the second inverted-V 255 structure. Two longer clusters of events are centered near 490 and 610 seconds. The per-bagel 256 set sizes are shown in the table inset in the upper left of Figure 5, showing that the majority 257 of the events were at 260 and 630 eV, with 12 and 23 events, respectively. The 820 and 1050 258 eV bagels had no qualifying timeslices. 259

The large number of correlation events inspires a search for statistical patterns in the 260 data which would not have been discernible in previous experiments due to their small number 261 of events. The first step is to decompose the events into resistive and reactive components. The 262 resistive component consists of particles which oscillate in phase or 180 degrees out of phase 263 with the electric field of the Langmuir wave, and the reactive component particles oscillate 90 or 270 degrees out of phase (Kletzing & Muschietti, 2006). A strong resistive component is 265 an indicator of wave-particle energy exchange, leading to wave growth or damping, while the 266 reactive phase is associated with particle trapping. The summation of these two components 267 will tend to have a sinusoidal form when either component has a significant value, and it is 268 this form that the event-identification method focused on. 269

To examine the correlator data in light of the theory, correlator z-scores at each timeslice were fit, using a nonlinear least squares analysis, to a quadrature function vs. bin number p,

$$-I\sin\left((p-p_0)*\frac{\pi}{8}\right) - Q\cos\left((p-p_0)*\frac{\pi}{8}\right),$$
(7)

adjusting I and Q to provide the best fit to the observed z-scores, starting from I = Q = 0.5. The negative signs and p_0 are determined by calibration data relating the electric-field phase to the bins. The coefficients I and Q in principle represent the strengths of the In-phase and Quadrature signals, i.e. the resistive and reactive linear components. An example set of fits



Figure 4. An example of data used to hand-screen events. (a) A spectrogram of 'bagel' detector data; (b) the logarithm of raw total counts for the selected bagel detector, with the correlator lock/no-lock status displayed at the bottom as an X for *lock*; (c) an HFE spectrogram covering the relevant frequency range for Langmuir waves, with overlays for the correlator lock frequency (solid white trace) and plasma frequency identified by inspection (dashed black trace); (d) HFE electric field waveform for a much smaller time interval: $\pm 4000 \ \mu$ s around the center time; (e) HFE waveform for an even more expanded time interval: $\pm 100 \ \mu$ s. This event was discarded because it shows a weak and noisy Langmuir cutoff, strong periodic interference at ~15 μ s cadence, and a fluctuating lock frequency with sporadic *lock* status.



Figure 5. An overview of the final event set S vs. time for the whole flight, with the per-bagel totals in the table inset top left. Individual events are displayed as x-marks on lines corresponding to bagels on the left vertical axis. Overlaid on this is a histogram of events vs. time, corresponding to event/bin counts on the right vertical axis.

are shown in the line plots on the right of Figure 2, in which the fit has been applied to the 276 middle (t = 50 ms) timeslice, with bins on the y-axis (aligned to the left σ plot)—in both 277 line plots, the black line shows the σ values for the center timeslice. In the left plot, the 278 solid blue line shows a normalized (unity amplitude) wave in-phase with the calibrated electric 279 field, corresponding to I, the resistive component, and the dashed red line is the reactive Q280 component. The right plot shows fitted forms, with blue and red the fitted, separate I and Q281 contributions, and the dot-dashed magenta line their sum (i.e. the actual fit function). For 282 this timeslice, the r^2 goodness-of-fit is ~ 0.906, showing a reasonable fit, and with $I \sim 0.3$ and 283 $Q \sim -3.1$ this event appears to be dominated by the reactive component. While many of the 284 selected events show r^2 values below 0.5 (r below 0.707), examination of the fitted data shows 285 that they are successful at reproducing the sinusoidal characteristics with correct phases in the 286 correlator data, even when r^2 is as low as 0.2 (r as low as 0.45), which covers the vast majority 287 of the selected events (51 of the 57). Plots of the event fits are included in the Supporting 288 Information. 289

Figure 6 shows results of fitting I and Q functions for the 57 selected events, expressed as 290 the log of the magnitude of the resistive-to-reactive coefficient ratio, $\log_{10} |I/Q|$, with vertical 291 lines drawn to divide the events by bagel. Some basic statistics of I/Q and |I/Q| are inset 292 on the lower right, and the events are color-coded by their r^2 goodness-of-fit value. Individual 293 plots of the fits are provided in the Supporting Information. There is significant scatter, but the 294 majority of events (42 out of 57), have $\log_{10} |I/Q| > 0$, implying they tend to be dominantly 295 resistive, although most events have both a resistive and reactive component. This result 296 could arise partly from experimental bias: the reactive component is bipolar, implying that 297 if the energy resolution is insufficient the correlation is more easily washed out than that of 298 the resistive component, which is unipolar in energy. However, this bias only applies if the 299 energy resolution is insufficient. The observation of both reactive and resistive components 300 in multiple events suggests that the energy resolution is adequate for many cases, as is also 301



Figure 6. The ratio of the resistive to the reactive component $(\log_{10} |I/Q|)$ of each of 57 selected waveparticle correlation events, with vertical lines dividing the events by bagel. Statistics of I/Q and |I/Q| are inset in the lower right, and events are color-coded by their r^2 goodness-of-fit value. Note that the majority of events show more resistive power in the fit than reactive. Individual plots of the fits are provided in the Supporting Information.

known from previous correlator experiments (Kletzing et al., 2005) and suggested by theory 302 (Kletzing & Muschietti, 2006). The tendency of the resistive component to be dominant may 303 appear somewhat counter-intuitive, since one might expect the wave growth/damping stage to 304 be of short duration compared to the static case in which electrons are trapped in the wave but 305 not significantly exchanging energy. However, as shown and discussed below, the majority of 306 selected cases correspond to time variations of the causative electron beam, conditions which 307 might be associated with wave growth/damping and hence energy exchange between electrons 308 and the waves. 309

Figure 7 shows four example events. In each case, for context, the top panel shows an electron spectrogram from the EEPAA, with high energy resolution and range but poor time resolution, covering one second centered on the event. The lower three panels show 41 ms of data around the event: a spectrogram of the total counts of each bagel vs. bagel energy, the values of I and Q, and a color plot of σ versus wave phase and time in the same format as the left panel of Figure 2.

The top left panels of Figure 7 show data from an event at 607.203 s after launch, involving 316 260 eV electrons. There is evidence of higher-energy (>500 eV) beams in the EEPAA data 317 leading up to the event, but no significant fluxes at higher energy during the event. The bagel 318 spectrogram reveals sporadic, short-duration (2-5 ms) 260 eV beams. The selected event occurs 319 in the middle of the time frame, at 607.203 s, identifiable as a column of high σ values with 320 strong negative (red) values at phases <180 degrees, and strong positive (blue) values at phases 321 >180 degrees. This phase pattern corresponds to a fit in which I < 0, called a 'negative-I' 322 event. The negative-I event occurs on the leading edge of the short-duration beam. A bit 323 further on in the record, associated with the short-duration beam at 607.219 s, there occurs a 324 negative-I event on the leading edge of the beam, and a positive-I event on the trailing edge. 325



Figure 7. Example wave-particle correlation events from among 57 identified by a combination of automated and manual screening. Each set of four panels includes: (a) one second spectrogram of EEPAA data showing context for the event; (b) 40-ms spectrogram of bagel data; (c) line plots of reactive/resistive fit values; and (d) σ versus wave phase versus time.

Density Gradient (top) and Wave Power Gradient vs. I/Q Power



Figure 8. Scatter plots of bagel count (i.e. resonant electron density) gradient vs. I (reactive, left column) and Q (resistive) fit values, with color as the r^2 goodness-of-fit value. Note the clear relation in the left plot, and the lack of such on the right.

This pattern persists in the other three examples of Figure 7. The top right panels show 326 two closely spaced examples, at 718.274 s and also involving 260 eV electrons, occurring on the 327 trailing edge of a beamlet and the leading edge of a subsequent one. Again the example on the 328 trailing edge is positive-I, and that on the leading edge is negative-I type. The examples in the 329 bottom panels differ in that they span multiple timeslices rather than a single timeslice as in 330 the top panels. The bottom left panel, at 596.234 s and involving 320 eV electrons, shows a 331 positive-I event spanning a couple of timeslices that occurs in a region in which the 320 eV and 332 820 eV beam densities are dropping off. The bottom right panel, at 652.845 s and involving 333 630 eV electrons, shows two events, each spanning a range of timeslices at the front and back 334 edge of a short-duration beam which is most prominent at 820 eV, but shows some structure 335 at lower energies. Following the pattern, the event on the leading edge where the beam density 336 increases is negative-I type, and that on the trailing edge where the beam density decreases is 337 positive-I type. 338

To explore this relationship further, Figure 8 shows scatter plots of the temporal gradient 339 of the electron beam flux $(\nabla_t n_B)$ inferred from the bagel detectors versus the I and Q fit values. 340 The points are colored according to their r^2 goodness-of-fit value from the reactive/resistive 341 fitting. The left panel demonstrates from this statistical approach the correlation between the 342 I value and the beam flux gradient which was illustrated by multiple examples in Figure 7. A 343 clear trend is evident whereby negative I values correspond to positive beam flux gradients, and 344 vice versa. A linear regression to these points returns a t-statistic value of -6.97 with a p-value 345 of 4×10^{-9} , implying that the null hypothesis—that there is no relation between I and $\nabla_t n_B$ — 346 should be rejected. The right panel of Figure 8 shows that there is no pattern evident between 347 the Q fit value and $\nabla_t n_B$. Linear regression of the right panel has a small t-statistic and 348 p-value ≥ 0.15 , suggesting that the relationship between I and $\nabla_t n_B$ is significant. Table 1 349 summarizes additional statistical tests performed on the data, showing that a two-sample 350 Kolmogorov-Smirnov test finds a significant difference between the $\nabla_t n_B < 0$ and $\nabla_t n_B > 0$ 351 distributions of I, and that in linear cross-correlations I and $\nabla_t n_B$ are significantly correlated. 352

Correlation might be expected between the I and Q fit values and the gradient in wave 353 power for cases in which the resistive component dominates, because such correlations imply 354 exchange of energy between electrons and waves. Such a correlation should be observable if 355 the wave power at a given frequency is dominated by the wave mode that is interacting with 356 the energy of electrons found to be correlated; however, the expected correlation would be on 357 timescales less than the 1 ms interval over which the correlation is detected. The wave data 358 appear inadequate to address this issue unless the gradient is larger than tens of dB over 0.1 359 ms, and gradients of this amplitude were not observed during any of the events. 360

361

Table 1. Statistical tests on the Figure 8 scatter pic	Table 1.	Statistical	$tests^a$ of	on the	Figure	8	scatter	plo
---	----------	-------------	--------------	--------	--------	---	---------	-----

	I - $\nabla_t n_B$	Q - $\nabla_t n_B$	
Linear Regression Kolmogorov-Smirnov Test Cross-Correlation	$\begin{array}{ccc} {\rm s} & {\rm p^b} \\ -6.97 & 4.0 \times 10^{-9} \\ 1^{\rm c} & 2.8 \times 10^{-8} \\ -0.68 & 7.8 \times 10^{-9} \end{array}$	$\begin{array}{cccc} & & p \\ 2.26 & 0.027 \\ 0 & 0.23 \\ 0.24 & 0.062 \end{array}$	7

^a These statistics evaluate relations from I and Q to $\nabla_t n_B$. The 's' columns contain the significant output of the given test: from top to bottom, the t-statistic, null-hypothesis rejection, and correlation coefficient.

^b In all tests, note the extremely low p-value of the $I-\nabla_t n_B$ relation compared to the $Q-\nabla_t n_B$ p-value, implying that this column's comparisons yield strong statistical results.

^c The Kolmogorov-Smirnov test's result shows only that the null hypothesis (that the distributions are the same) is accepted or rejected. Here, note that it is rejected for the $I - \nabla_t n_B$ relation.

362 4 Discussion

During the CHARM-II flight, the correlator detected 57 wave-particle correlation events 363 deemed significant after application of the automatic algorithm followed by manual inspection 364 and screening. An analysis of the reactive and resistive components, I and Q, of these events 365 reveals a correlation between a positive electron beam flux temporal gradient $\nabla_t n_B$ at a given 366 bagel energy level, as in the case of an electron beam appearing at that energy level, and a 367 negative value of I for the coincident wave-particle correlation event, and a similar relation 368 between a negative $\nabla_t n_B$ and positive I. Careful consideration of the correlator calibration 369 confirms that the positive half of the electric-field waveform corresponds to a field pointing 370 towards the bagels, and thus electrons being accelerated away from them. Thus, the observed 371 relation is consistent with energy going from the beam to the wave field during a beam density 372 increase, and the inverse for a density decrease. The lack of evidence for a relation between the 373 HFE power and I and Q is curious, given prior observations by Kletzing et al. (2005) of such a 374 relation in a case study of data from the RACE sounding rocket. It is possible that the extreme 375 wave power during the majority of the 57 selected events may mask such an effect, particularly 376 given that the amplitude modulation typical of the bursty Langmuir waves observed in the 377 RACE experiment is not prevalent in the CHARM-II HFE data. 378

It is plausible that time-of-flight effects on a time-varying electron beam may lead to a 379 situation in which Langmuir waves experience spikes in growth or damping like those seen in 380 Figure 7. Given a beam population 'turned on' promptly at some altitude above the rocket, and 381 a 'warm' background population of degraded secondary electrons associated with the beam, as 382 depicted in Figure 9, the highest-energy particles arriving earliest at the rocket location may 383 create a positive-slope region in the reduced parallel distribution function. This will produce 384 wave growth, characterized by transfer of energy from electrons to waves and detected by the 385 correlator as a negative-I type wave-particle correlation. In contrast, when the beam 'turns 386 off,' it is also possible, depending on the relative energies and densities of the two electron 387 populations, that an enhanced negative slope will appear in the reduced distribution function 388 as the higher-energy particles disappear first, leading to enhanced damping of the waves due 389

Reduced Distribution Function



Figure 9. A cartoon showing an explanation for the relation seen between I and $\nabla_t n_B$. As a beam appears, the high-energy particles are the earliest to arrive, leading to an exaggerated positive slope and wave growth (left). The remainder of the particles and distribution relaxation then yield a plateau in middle times. Finally, when the beam turns off at the source, the high-energy particles are the first to disappear, and in the right configuration may yield an exaggerated negative slope, enhancing wave damping (right).

to interactions with electrons in the narrow energy region, and resulting in a wave-particle correlation event of the positive-I type. The presence of short-lived beam features in the topleft and bottom-right plots of Figure 7, and their temporal relation to the nearby correlator events, provide evidence suggesting this mechanism.

An extreme case of such time-of-flight effects which is easy to perceive physically is the 394 dispersive electron beam which has been observed frequently, e.g. by Kletzing et al. (2016). 395 This type of dispersive beam, in which higher energies arrive before lower energies, is the 396 normal pattern for parallel electron beams in so-called Alfvénic aurora, in which the beams 397 are accelerated by Alfvén waves at altitudes well above the rocket and undergo dispersion 398 as they propagate to lower altitudes (Kletzing & Hu, 2001; Chen et al., 2005). The beam energy decreases with time from left to right, and the energy range of an appropriate fixed-400 energy particle detector will shift from lying below the peak energy of the beam to lying 401 above the peak energy. In the former case the detected energy corresponds to the condition 402 $df/dv_{\parallel} > 0$, which yields Langmuir wave growth, and the latter case corresponds to $df/dv_{\parallel} < 0$ 403 which damps Langmuir waves. Under the former condition one expects waves resonant to the 404 detector's energy to be growing, extracting energy from the beam, which would correspond 405 to the negative values of the in-phase component of the electron-electric field correlation. 406 Under the latter condition, the opposite energy flow would be expected, corresponding to wave 407 damping at the detector energy. The expected signature in the phase of the electron bunching 408 is exactly as observed. 409

Langmuir wave growth during an increase in the number of electrons at or near the resonant energy is generally expected because of the resultant instability, whether due to a beam moving into an energy range or simply appearing at that energy. While subsequent damping is also expected, an impulsive enhancement of damping concurrent with the beam's disappearance, is, on the other hand, not an immediately obvious causal relationship. The easiest way to try to visualize and confirm the effect is by using a numerical simulation.

To this end, numerical calculations have been performed using a a test-particle simulation tool described in Appendix A. The basis for this calculation is Liouville's theorem

which allows individual test particles to stand for portions of the electron phase space distribu-418 tion. A library of particle trajectories has been created by launching test particles into a model 419 converging magnetic field, and allowing 5000 km of simulated flight. For computational effi-420 ciency the calculation proceeds backwards, with test particles representing the full distribution 421 function at the low-altitude end traced backwards in time and space to find out which region of 422 phase and physical space they connect to at the top end. It is then possible to model arbitrary 423 input electron distributions at the top end, and time variations in those, and determine the electron distribution as a function of time at the bottom end. As a final step in the numerical 425 calculation, the reduced distribution is calculated from the full distribution function, and the 426 Langmuir wave growth or damping rate is calculated from the reduced distribution function. 427 Details of this multi-step numerical tool are given in Appendix Appendix A. 428

For this paper, the simulation was run with ionospheric background parameters $f_{pe} = 400$ kHz and $T_{iono} = 2000$ K, secondary background $T_{bg} = 2 \times 10^5$ K with $n_{bg} \simeq 1.07 \times 10^6$ m⁻³, and beam parameters $T_{beam} = T_{bg}/5$ and $n_{beam} = n_{bg}/50$ with a velocity shift $\delta(t) = 400$ m/s.

Figure 10 shows the final results of the simulation of Langmuir wave growth/damping, 432 resulting from a beam 5000 km above the observation point that was turned on promptly at 433 t = 5 s and off promptly at t = 10 s. The top panel shows the ratio of the beam density to 434 the background density at the top end, as a fraction of time as imposed in the simulation. 435 The middle panels show the maximum Langmuir wave growth rate γ , in units of inverse 436 seconds, as a function of time—and as a color scale (red=growth, blue=damping)—for two 437 selected intervals, on the left starting two seconds after the beam was turned on, and the 438 right starting two seconds after it was turned off. The lower panels show growth/damping 439 rate on a symmetric-logarithmic color scale as a function of wavenumber k and time for the 440 441 same two selected time intervals. All these calculations correspond to a selected frequency range just above the plasma frequency, with parameter ω_t (described in Appendix Appendix 442 A) ranging from approximately ω_p to $1.02\omega_p$, corresponding to k = 0.01 to 1. The lower limit 443 of k is determined by the resonance condition $kv_{\parallel} = \omega_p$. For the theory to be viable v_{\parallel} should 444 be kept far enough below c that any relativistic corrections remain small—in our simulation, 445 where $\omega_p = 400$ kHz, this sets a lower bound on k of $10^{-2}m^{-1}$. 446

Significantly, Figure 10 shows both a growth rate spike during the beam arrival and a damping enhancement during beam departure as predicted by the qualitative argument above, thus supporting the mechanism advanced for the relationship between the direction of energy flow in the wave-particle correlation and the gradient in the beam density. This result matches qualitatively the pattern observed in the phase of the in-phase, resistive component (I) of the wave-particle correlations during positive and negative gradients in beam flux.

The bottom panel of Figure 10 shows that the strongest growth and damping are associated with the long-wavelength modes; this is generally expected, as shorter-wavelength modes are more heavily damped. Growth at the long wavelengths is associated with the earliestarriving, higher-energy particles, with the later lower-energy arrivals exciting some growth at shorter wavelengths.

The overall time frame of the growth and damping peaks are of order 100 ms, which is 458 significantly longer than most of the observed wave-particle correlation events; however, there 459 are suggestions of short-timescale structures in the simulation events, for which the growth 460 and damping rates exhibit multiple shorter-timescale peaks. The timing of these fine features, 461 however, has been found to shift significantly based on parameters of the simulation, such as 462 shifting the beam up and down in energy, or removing parts of the beam distribution function, 463 although the overall result showing enhanced growth and damping is robust (Dombrowski, 464 2016). These tests suggest that the fine structure seen in Figure 10 may depend on aspects of 465 the simulation system. 466

Another variable is the fact that the instrumentation is in motion. Given that the payload's total velocity ranges from 1300 to 2600 m/s during the flight period of interest and the



Growth Rates, Δt_s =0.001 s, Δt_p =0.01 s

Figure 10. Results from the simulation: Langmuir wave growth rate γ , versus k and time (on the horizontal axis), calculated for ω_t from approximately ω_p to $1.02\omega_p$. The top panel shows the n_{beam}/n_{bg} at the top, while the two columns are zoomed into the times at which the bulk of the particles arrive during beam turn-on (left) or depart with beam shutoff (right), after an approximate 2-second travel time—for the fastest particles—across the length of the simulation. The top growth-rate panels show γ on the vertical axis, as well as in color scale (blue is negative, red positive), and both a growth rate spike during the beam arrival and a damping enhancement during beam departure are clearly visible, qualitatively matching the the pattern observed in the data. In the lower panels, the color scale is still γ , and the vertical axis is the wavenumber k. The lower limit of the plot is determined by the resonance condition $kv_{\parallel} = \omega_p$, with v_{\parallel} kept to where any relativistic correction is small.

dynamic, non-uniform nature of the aurora, it is plausible that the short duration of detected
events is due to payload transitioning in and out of plasma structures with limited spatial extent. To deconvolve the effects of payload motion on correlation measurements, synchronized,
multi-point observations would be required.

Further investigation using more sophisticated tools is required to show whether the
mechanism described here can produce temporospatial structuring in wave-particle correlations
on the observed few-ms timescales. The numerical simulation described here does, however,
prove the overall plausibility of the mechanism.

477 5 Conclusions

The CHARM-II sounding rocket carried a Langmuir wave receiver, eight large-geometric 478 479 factor upward-looking electron detectors tuned to tightly spaced energies, and a wave particle correlator to an apogee altitude of 802 km in substorm aurora, returning data above 600 km 480 from approximately 500 s after launch. Applying both automated and manual event selec-481 tion, 57 timeslices containing wave-particle correlation events were identified as statistically 482 significant and analyzed. Breakdown of the phase correlation data into resistive and reactive 483 components revealed a striking relationship between electron beam dynamics and the nature of the wave-particle correlation: whenever the beam flux at the measured electron energy was 485 increasing with time, the phase of the resistive component of the electron bunching implied 486 energy transfer from the beam particles to the wave field, and when the electron beam flux was 487 decreasing, the reverse occurred. This pattern was repeated for all events, and was particularly clear in several events, including the largest-amplitude event investigated by Kletzing et al. 489 (2016).490

Two related theories to explain this observation have been explored, one invoking the 491 changing nature of the interactions of the electrons with a given Langmuir wave as the beam 492 energy decreases, as typically occurs due to dispersion of an auroral electron beam accelerated 493 several thousand kilometers above the interaction location; and the other invoking detailed 494 features of the electron distribution function at ionospheric altitudes, arising when the electron 495 beam is modulated at higher altitudes. A magneto-kinetic test-particle numerical simulation 496 confirmed that for an electron beam which causes an impulsive increase in wave growth upon 497 its appearance, its disappearance will be accompanied by an impulsive enhancement of wave 498 damping within the same frequency range. The results therefore agree qualitatively with 499 the experimental data from the CHARM-II rocket, though an exactly simulated quantitative 500 representation has not yet been achieved. 501

502 Appendix A Simulation

The aim is to simulate a minimal-complexity environment that is sufficient to probe 503 whether an electron beam with reasonable characteristics, and which shows significant Lang-504 muir wave growth upon its appearance, also shows enhanced wave damping as it disappears. 505 This goal requires calculating the short-timescale behavior of the wave growth rate as the elec-506 tron distribution function evolves in time due to the temporal turning on and off of the beam, 507 combined with travel-time effects in the converging magnetic field geometry. These factors 508 compel the use of a complex, flexible—and computationally intensive—test particle simulation 509 system. Its application to this case shall use simple, gyrotropic magneto-kinetic parameters, 510 with no inter-particle interaction or wave-particle scattering. 511

Following a numerical analog to the analytical method of Cairns (1987), Liouville's equation governs the evolution of a distribution function over time, and with no wave-particle or inter-particle scattering, one can write

$$f(\bar{x}, \bar{v}, t) = f(\bar{x}', \bar{v}', t'),$$
 (A1)

⁵¹⁵ i.e. that the value of the distribution function at a source phase-space region (\bar{x}', \bar{v}') at time ⁵¹⁶ t' is the same for the related region (\bar{x}, \bar{v}) at time t. The test-particle simulation is used to ⁵¹⁷ relate the primed and unprimed regions, by creating a lookup table of particle travel times ⁵¹⁸ $\mathbb{T}(\mathsf{E}, \alpha')$ for a range of source energies E and pitch angles α . These test particles are then ⁵¹⁹ treated as centers of regions in phase space, and are used to 'carry', in time \mathbb{T} , values of the ⁵²⁰ source distribution function down to a corresponding region (E, α) at the observation point.

In this analysis, z is taken to be a positive-upward, field-aligned coordinate, with z = 0521 corresponding to the beam generation altitude. To simulate only those particles which will 522 arrive at the 'detection point' at (x = y = 0, z = -5000 km) requires using a deterministic 523 (i.e. time-reversible) simulation method: we watch for when test particles originating at the 524 detection point with an upward velocity cross a target plane at z = 0. The velocities can then 525 be reversed for the later downgoing analysis. The 'Boris Method' is used—a standard, time-526 reversible particle pusher (Boris, 1970; Birdsall & Langdon, 2005). This method separates 527 the effects of the electric and magnetic forces, dividing them into a half-impulse from any 528 background electric field, followed by a rotation according to the magnetic field, and then 529 another electric half-impulse. 530

⁵³¹ Careful testing of energy conservation led to setting a unitless timestep of 0.01. The base ⁵³² of the time system is the electron cyclotron period, and so this is equivalent to each timestep ⁵³³ moving each particle a hundredth of an orbit. For the input parameters used, this yielded a ⁵³⁴ worst-case energy loss of 0.06% over the full length of the simulation.

To allow a realistic amount of time/space for mixing of particles of different energies and pitch angles, a distance of 5000 km is used, corresponding to the distance from the bottom of the electron acceleration region to the ionospheric detection point. The background electric field is assumed zero, and the magnetic field is rotationally symmetric around x = y = 0, defined as

$$\bar{B} = -\frac{zr}{L^2}\hat{r} + \frac{1+z^2}{L^2}\hat{z} = -\frac{xz}{L^2}\hat{x} - \frac{yz}{L^2}\hat{y} + \frac{1+z^2}{L^2}\hat{z},\tag{A2}$$

with L a scaling variable determined by our desired mirror ratio and target distance. For the following simulation, the mirror ratio was set to 5.

In order to fully cover the range of energies detected by the wave-particle correlator, 542 particles were launched at 31 energies linearly spaced in velocity, with energies ranging from 543 25 to 1225 eV. At each energy, particles were launched at each of 41 pitch angles, with an 544 angular spacing of $3\pi/256$. Because of the rotational symmetry of the simulation imposed by 545 gyromotion, it is only necessary to launch particles at one azimuthal angle—the results can 546 then be rotated to fill a velocity-space hemisphere at the detection point. To ensure that our 547 hemispherical segments sweep out constant solid angles $\Delta\Omega$ —herein set to 0.001 steradians— 548 requires that the azimuthal launch angles ϕ be related to by their separation $\Delta \phi$ to the pitch angle separation $\Delta \alpha$ such that $\Delta \phi = \frac{\Delta \Omega}{\sin(\alpha) \Delta \alpha}$. We determine and use the integer number of 549 550 angles ϕ which most closely achieve this relation. 551

The simulation code was implemented in MATLAB, manually fragmented into 64 shards, and run in approximately five weeks on the Dartmouth Discovery cluster. Due to the finite timesteps, which are unlikely to land precisely on z = 0, interpolation was required to find exact crossing parameters. To enable this, the final 1000 timesteps for each particle were saved (with the very last step having z > 0), and gyro-orbit equations were fit to these, from which accurate final z = 0 position, velocity, and travel times come forthwith. All unitless values were then interpreted via inter-defined base values: $B_0 = 50$ microtesla, $t_0 \approx 714$ ns, $r_0 \approx 0.337$ m, and $v_0 = 0.00989$ c, corresponding to 25 eV.

Figure A1 shows some basic diagnostics of the output of this simulation, with expected trends compared to total energy and launch pitch angle. Once the table $\mathbb{T}(\mathsf{E}, \alpha)$ has been created, it becomes possible to reverse the velocities and launch distributions downward, towards the simulated 'detector'.



Figure A1. Various diagnostics of the input characteristics and resultant travel times in the testparticle simulation. All time axes are travel time along the length of the simulation. Energy and pitch angle in the lower row are the launch values at the detection point (high magnetic field).

The beam distribution is imposed at the top (lowest B) on total velocity, |v| (i.e. we assume $T_{\parallel} = T_{\perp}$ and a flat distribution across pitch-angle space), and sampled at the test velocities used in S. The beam is 'turned on' (or off) instantly at a specified time by adding and removing an additional component to the particle distribution sum. The environment is assumed to be both homogeneous and large enough that any generated region of velocity-space at the top will detected at the bottom. Thus, the x and y positions of particles in the simulation are neglected.

For purposes of calculating growth rate, the distribution at the rocket altitude is pre-571 sumed composed of three components. First, there is a cold background, f_{iono} , with a 2000 K 572 temperature which is taken to be uniform and constant, unaffected by the particle tracing, and 573 mainly plays the role of contributing the largest part of the plasma density, hence essentially 574 determining the plasma frequency. A secondary background distribution serves as the main 575 background to the beam for growth-rate calculation purposes. It is a localized population, 576 distinct from the much colder, higher-density ionospheric background, and is composed of a 577 population of degraded beam particles with a higher density and temperature than the beam 578 population. It is defined as an isotropic Maxwellian, 579

$$f(|v|) = n_e \left(2\pi\right)^{3/2} v_{th}^3 \,\mathsf{e}^{-\left(\frac{v}{2v_{th}}\right)^2} \Rightarrow f_{bg} = n_{bg} \left(\frac{m_e}{2\pi k T_{bg}}\right)^{3/2} \mathsf{e}^{-\frac{m_e v^2}{2k T_{bg}}},\tag{A3}$$

where, for a given input temperature, the electron number density n_{bg} is interpolated from a table of values used by Lotko and Maggs (1981). The third population is the beam, which is also originally Maxwellian, isotropic, and homogeneous in velocity space at the top of the simulation region, except that its parameters can be considered time-varying $T_{beam}(t)$, $n_{beam}(t)$, and a velocity shift $\delta(t)$:

$$f_{beam}(|v|,t) = n_{beam}(t) \left(\frac{m_e}{2\pi k T_{beam}(t)}\right)^{3/2} e^{-\frac{m_e(v-\delta(t))^2}{2k T_{beam}(t)}}.$$
 (A4)

In practice, $T_{beam}(t)$ and $n_{beam}(t)$ are set as fractions of the secondary background values. The final distribution is the sum of these,

$$f(v,t) = f(v_{\parallel}, v_{\perp}, \phi, t) = f_{iono} + f_{bg} + f_{beam}(t).$$
 (A5)

To dimensionally reduce these towards a parallel distribution function $f(v_{\parallel})$, the first step is to sum over the azimuthal angles. This is not a simple sum: as these are finite cells in velocity-space, each angular 'wedge' must be weighted by its accompanying $\Delta \phi_i$, i.e.

$$f(v_{\parallel}, v_{\perp}, t) = \sum_{i \forall \phi} f(v_{\parallel}, v_{\perp}, \phi_i, t) \Delta \phi_i,$$
(A6)

where $\Delta \phi_i$ is set by the pitch angle, as in the hemispheric interpolation.

Time variations have been neglected until now because the hemispheric interpolation 591 introduces no time dependence, and so the azimuthal sum has none either. Simulating the 592 distribution function from the electrons being emitted effectively continuously from a source 593 region requires setting a Δt_D period over which the detector bins incoming particles, and 594 a Δt_s period between source distribution 'launches'. To achieve something approaching the 595 appearance of a continuous source, Δt_D should be at least $10\Delta t_S$. The next step is interpolating 596 and reducing away a dimension from the test-particle simulation, while taking travel times into consideration. This is simply a set limitation at detector timeslice τ , such that the particles 598 we consider are, henceforth, in the set \mathbb{J}_{τ} of particles whose launch time t_0 and travel time t_T 599 fulfill $t_0 + t_T \leq \tau$ and $> \tau - t_D$. There is also an implicit sum here as the detector integration 600 is accounted for, which requires its own weighting value for f, simply the ratio of the launch 601 time and the integration time, i.e. 602

$$f(v_{\parallel}, v_{\perp}, \tau) = \frac{\Delta t_S}{\Delta t_D} f(\mathbb{J}_{\tau}).$$
(A7)

Therefore, the second step in obtaining a one-dimensional reduced distribution function, after summing over the azimuthal angles, is to perform a time-limited, time-weighted sum over the perpendicular velocities, following the standard Landau theory for parallel propagation (Ergun et al., 1993, e.g.). This is slightly complicated by the large number of unique v_{\parallel} values. Defining 'center points' in v_{\parallel} as simply the points along the $v_{\perp} = 0$ axis, for each of these $v_{\parallel,i}$, all values of f with v_{\parallel} in the range $\mu_{\parallel} = \left\{ \frac{v_{\parallel,i-1} + v_{\parallel,i}}{2}, \frac{v_{\parallel,i} + v_{\parallel,i+1}}{2} \right\}$ are summed over v_{\perp} using a modification of the trapezoidal rule,

$$f(v_{\parallel,i},\tau) = \sum_{j \forall v_{\perp}; \mu_{\parallel}} (v_{\perp,j+1} - v_{\perp,j}) \frac{f(\mu_{\parallel}, v_{\perp,j}, \tau) + f(\mu_{\parallel}, v_{\perp,j+1}, \tau)}{2} v_{\perp,j},$$
(A8)

where the factor $v_{\perp,j}$ is the phase-space cell weighting. Figure A2 shows a color plot of the reduced distribution function vs. v_{\parallel} and time, as well as several timeslices as the distribution evolves through the beam-arrival phase.

Finally, Langmuir wave growth rate can be calculated from the time-integrated onedimensional reduced distribution function in v_{\parallel} . For a given cold ionospheric background plasma frequency ω_p , wave vector $\mathbf{k} = k_{\parallel} \mathbf{B}/B$, and test wave frequency ω_t given by the dispersion relation, the growth rate is

$$\gamma(f(v_{\parallel}), k_{\parallel}, \omega_p, \omega_t, \tau) = \left(\frac{d\epsilon}{d\omega}\right)^{-1} \operatorname{Sign}[k] \frac{\pi \omega_p^2}{k^2 n_e} \left[\frac{\partial f(v_{\parallel}, \tau)}{\partial v_{\parallel}}\right]_{kv_{\parallel} = \omega_t},\tag{A9}$$

where ϵ is the dielectric function, approximated as $1 - \frac{\omega_p^2}{\omega_t^2}$ for cold plasma. The derivative $\partial f(v_{\parallel}, \tau) / \partial v_{\parallel}$ is calculated at a test velocity related to the beam parameters, specifically the closest v_{\parallel} value to the resonance condition, ω_t/k .

Ideally, Δt_D should match the correlator's time resolution, 1 ms, and also allow the simu-620 lation to 'settle' for a long enough time between source changes that even the slowest particles 621 reach the detector, approximately 14 s per change. However, given the above guideline that 622 $\Delta t_S \leq \Delta t_D/10$, this would require storage of prohibitive numbers of time-overlapping distri-623 butions, so some sort of approximation is necessary. Figure A1 shows that the majority of 624 particles arrive within 5 seconds, so taking that as the settling time is a reasonable approxima-625 tion. The calculations shown in the paper use Δt_D of order 10-50 ms; short test calculations 626 with $\Delta t_D \sim 1$ ms showed no obvious qualitative differences, but were impractical for the 627 simulations in Figures 10 to A2. As an additional approximation, k and ω_t are determined 628 via the warm plasma dispersion relation, $\omega_t = \omega_p + \frac{3}{2}k^2 v_{th}^2 / \omega_p$, where v_{th} is the background 629 ionospheric thermal speed $\sqrt{3kT_{iono}/m_e}$. With these approximations Langmuir wave growth 630 rates are calculated as a function of frequency, wavenumber and time at the rocket altitude, 631 for postulated time varying beams injected at the top of the simulation region, revealing vari-632 ations in growth or damping of the waves associated with time-of-flight effects on the particle 633 distribution function at the rocket altitude. 634

635 Acknowledgments

The authors thank engineers David McGaw and Jeff Dolan for their work with the particle 636 correlator and its components. The NASA/NSROC team led by Ted Gass executed a successful 637 launch from Poker Flat Research Range. The authors thank Dr. Wayne Scales for providing 638 the original magneto-kinetic test-particle code on which this simulation system was based, 639 and Susan Schwartz and Bill Hamblen for assistance running the code on the Dartmouth 640 Discovery cluster. CHARM-II sounding rocket data are archived at https://dartmouth.app 641 .box.com/v/CHARM2-rocket-40025-HF-data, and simulation codes are available on GitHub: 642 https://github.com/dc-tpgr-sim. Work at Dartmouth College was supported by NASA 643 through grant NNX17AF92G; work at University of Iowa was supported by NASA through a 644 sub-award of the same grant; work at the University of Sydney was supported by its Special 645 Studies Program. 646



Figure A2. Reduced distribution function values. Top shows a color plot vs. v_{\parallel} and time, for the entire time span of the test. Below, six timeslices from the beam arrival period, showing the formation and disappearance of a positive slope (compare panels c, d, and f to Figure 9). Note that the component of low-energy electrons centered near $0.5 \times 10^7 \text{ ms}^{-1}$ is the tail of the background (bg) distribution. It has not yet arrived before beam turn-on, but is considered too small to significantly affect the results.

647 References

659

660

661

662

663

676

677

678

679

680

681

682

683

684

685

686

687

688

- Birdsall, C. K., & Langdon, A. B. (2005). Plasma physics via computer simulation (S. Cow ley, P. Stott, & H. Wilhelmsson, Eds.). New York: Taylor & Francis Group.
- Boehm, M., Paschmann, G., Clemmons, J., Höfner, H., Frenzel, R., Ertl, M., ... Lundin, R.
 (1994, Apr). The tesp electron spectrometer and correlator (f7) on freja. Space Sci. Rev., 70, 509-540.
- Boehm, M. H. (1987). Waves and static electric fields in the auroral acceleration region (Doctoral thesis). University of California, Berkeley, Berkeley, CA.
- Boris, J. P. (1970, Nov). Relativistic plasma simulation-optimization of a hybrid code. In J. P. Boris & R. A. Shanny (Eds.), *Proceedings of the conference on the numerical*
- simulation of plasmas (4th) (p. 3-67). Washington, DC: National Technical Information Service.
 - Cairns, I. H. (1987, Mar). The electron distribution function upstream of earth's bow shock. J. Geophys. Res., 92(A3), 2315-2327.
 - Chen, L. J., Kletzing, C. A., Hu, S., & Bounds, S. R. (2005, Oct). Auroral electron dispersion below inverted-v energies: Resonant deceleration and acceleration by alfvén waves. J. Geophys. Res., 110(A10), A10S13. doi: 10.1029/2005JA011168
- Dombrowski, M. P. (2016). Sounding-rocket studies of langmuir wave microphysics in the auroral ionosphere (Doctoral thesis, Dartmouth College). doi: 10.1349/ddlp.1956
- Dombrowski, M. P., LaBelle, J., Rowland, D. E., Pfaff, R. F., & Kletzing, C. A. (2012,
 Sep). Interpretation of vector electric field measurements of bursty Langmuir waves
 in the cusp. Journal of Geophysical Research (Space Physics), 117, 9209. doi: 10.1029/2012JA017741
- Ergun, R. E., Carlson, C. W., McFadden, J. P., & Clemmons, J. H. (1991, Jan). Langmuir wave growth and electron bunching: Results from a wave-particle correlator. J. Geophys. Res., 96(A1), 225-238.
- Ergun, R. E., Carlson, C. W., McFadden, J. P., TonThat, D. M., & Clemmons, J. H. (1991, Jul). Observation of electron bunching during landau growth and damping. J. Geophys. Res., 96(A7), 11,371-11,378.
 - Ergun, R. E., Delory, G. T., Klementis, E., Carlson, C. W., McFadden, J. P., Roth, I., & Temerin, M. (1993, Mar). Vlf wave growth from dispersive bursts of field-aligned electron fluxes. J. Geophys. Res., 98(A3), 3777-3787.
 - Ergun, R. E., Malaspina, D. M., Cairns, I. H., Goldman, M. V., Newman, D. L., Robinson, P. A., ... Kaiser, M. L. (2008, Aug). Eigeinmode structure in solar-wind langmuir waves. *Phys. Rev. Lett.*, 101(051101).
 - Ergun, R. E., McFadden, J. P., & Carlson, C. W. (1998). Wave-particle correlator instrument design. In *Measurement techniques in space plasmas: Particles* (Vol. Particles, p. 325-331). American Geophysical Union.
 - Filbert, P. C., & Kellogg, P. J. (1979, Apr). Electrostatic noise at the plasma frequency. J. Geophys. Res., 84(A4), 1369-1381.
 - Gough, M. P., Christiansen, P. J., & Wilhelm, K. (1990, Aug). Auroral beam-plasma interactions: Particle correlator investigations. J. Geophys. Res., 95(A8), 12,287-12,294.
- Graham, D. B., & Cairns, I. H. (2013, Sep). Constraints on the formation and structure
 of langmuir eigenmodes in the solar wind. *Phys. Rev. Lett.*, 111, 121101. Retrieved
 from http://link.aps.org/doi/10.1103/PhysRevLett.111.121101
 doi: 10.1103/
 PhysRevLett.111.121101
- Graham, D. B., & Cairns, I. H. (2014, Apr). Dynamical evidence for nonlinear langmuir
 wave processes in type iii solar radio bursts. J. Geophys. Res., 119(4), 2430-2457. doi:
 10.1002/2013JA019425
- Gurnett, D. A., Scarf, F. L., Kurth, W. S., Shaw, R. R., & Poynter, R. L. (1981, Sep). De termination of jupiter's electron density profile from plasma wave observations. J. Geo *phys. Res.*, 86(A10), 8199-8212.
- Kaeppler, S. R., Kletzing, C. A., Bounds, S. R., Gjerloev, J. W., Anderson, B. J., Korth,
- H., ... Heinselman, C. J. (2013). Current closure in the auroral ionosphere: Results from the auroral current and electrodynamics structure rocket mission. In A. Keiling,

702	E. Donovan, F. Bagenal, & T. Karlsson (Eds.), Auroral phenomenology and magne-
703	tospheric processes: Earth and other planets (pp. 183–192). American Geophysical
704	Union. Retrieved from https://agupubs.onlinelibrary.wiley.com/doi/abs/
705	10.1029/2011GM001177 doi: 10.1029/2011GM001177
706	Kintner, P. M., Bonnell, J., Powell, S., & Wahlund, J. E. (1995). First results from the freja
707	hf snapshot receiver. Geophys. Res. Lett., 22, 287.
708	Kletzing, C. A., Bounds, S. R., LaBelle, J., & Samara, M. (2005). Observation of the re-
709	active component of langmuir wave phase-bunched electrons. Geophys. Res. Lett.,
710	32(L05106). doi: $10.1029/2004$ GL021175
711	Kletzing, C. A., & Hu, S. (2001, Feb). Alfvén wave generated electron time dispersion. Geo-
712	phys. Res. Lett., 28(4), 693-696. doi: 10.1029/2000GL012179
713	Kletzing, C. A., LaBelle, J., Bounds, S. R., Dolan, J., Kaeppler, S. R., & Dombrowski, M. P.
714	(2016, Feb). Phase sorting wave-particle correlator. J. Geophys. Res. Space Physics,
715	122, 20969-2078.
716	Kletzing, C. A., & Muschietti, L. (2006). Phase correlation of electrons and langmuir waves.
717	In J. LaBelle & R. A. Treumann (Eds.), (Vol. 687, p. 313-337). Springer Berlin Heidel-
718	berg.
719	Langmuir, I. (1928, June). Oscillations in ionized gasses. Proceedings of the National
720	Academy of Sciences, 14(8), 627-637.
721	Lin, R. P., Potter, D. W., Gurnett, D. A., & Scarf, F. L. (1981). Energetic electrons and
722	plasma waves associated with a solar type iii radio burst. Astrophys. J., 251(1), 364.
723	Lotko, W., & Maggs, J. E. (1981, May). Amplification of electrostatic noise in cyclotron res-
724	onance with an adiabatic auroral beam. J. Geophys. Res., 86(A5), 3449-3458.
725	Malaspina, D. M., Cairns, I. H., & Ergun, R. E. (2012, Aug). Antenna radiation near the lo-
726	cal plasma frequency by langmuir wave eigenmodes. Astrophys. J., 755(45).
727	Malaspina, D. M., & Ergun, R. E. (2008, Dec). Observations of three-dimensional langmuir
728	wave structure. J. Geophys. Res., 113 (A12108).
729	McAdams, K. L. (1999). Sounding rocket based investigations of hf waves in the auroral
730	<i>ionosphere</i> (Doctoral thesis, Dartmouth College), doi: 10.1349/ddlp.2820
731	McAdams, K. L., Ergun, R. E., & LaBelle, J. (2000, Feb). Hf chirps: Eigenmode trapping in
732	density depletions. Geonhus. Res. Lett., 27(3), 321-324.
733	McAdams, K. L., LaBelle, J., Schuck, P. W., & Kintner, P. M. (1998, Aug). Phaze ii obser-
734	vations of lower hybrid burst structures occurring on density gradients. <i>Geophys. Res.</i>
735	Lett. 25(16), 3091-3094.
736	Muschietti, L., Roth, L. & Ergun, R. E. (1994). Interaction of langmuir wave packets with
737	streaming electrons: Phase-correlation aspects Phus Plasmas 1 1008
738	Samara, M. (2005). Sounding rocket based investigations of whister, upper hubrid and
730	lanamuir waves in the auroral ionosphere (Doctoral thesis Dartmouth College) doi:
740	10 1349/ddlp 81
740	Samara M & LaBelle J (2006a Sen) Lf/mf whistler mode dispersive signals observed
741	with rocket-borne instruments in the auroral downward current region <i>J. Geonbus</i>
742	while receive both constrained in the darotal downward current region: v . <i>Coopingo</i> : $R_{es} = 111(A \cap 305)$
743	Samara M & LaBelle I (2006b) Structured waves near the plasma frequency observed in
745	three auroral rocket flights Annales Geonhusicae 2/(11) 2011-2010 Retrieved from
745	http://www.ann-geophys.net/24/2911/2006/ doi: 10.5104/angeo- 24 -2011.2006
740	Samara M LaBelle J Kletzing C A & Bounds S R (2004 Nov) Rocket observations
740	of structured upper hybrid waves at fub - 2fce Geonhus Res Lett 21(22) Ratrioved
740	from http://dx doi $org/10 1020/2004/CL021043$
149	1011 101p.//ux.u01.01g/10.1020/20040L021040